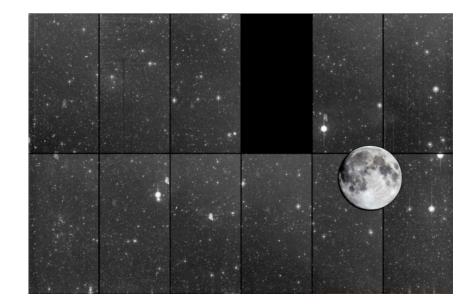
A new method to search for Supernova Progenitors in the PTF Archive

Frank Masci & the PTF Collaboration



The Palomar Transient Factory (PTF)

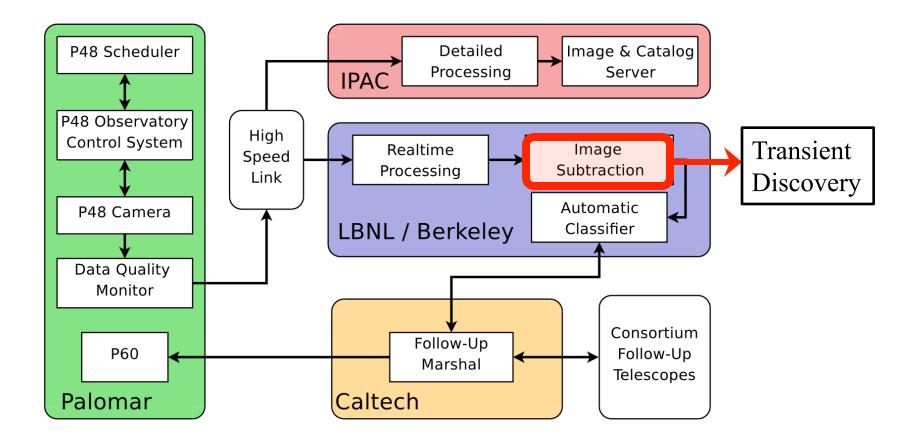
- Fully automated wide-field time-domain survey in optical
- Carried out on 48" Schmidt telescope on Mt Palomar
- 7.2 deg² FOV, 11 CCDs with ~ 92 megapixels, ~1"/pixel
- Image up to 2000 deg² per night (g and R filters) to 5σ depth of $R \sim 20.5$ with typical seeing $\sim 2''$
- Science operations began in March 2009
- >100 TB of image data and 10^{12} extracted sources **per year**
- Cadence: 1 min to 5 days: SNe, variable stars, asteroids...





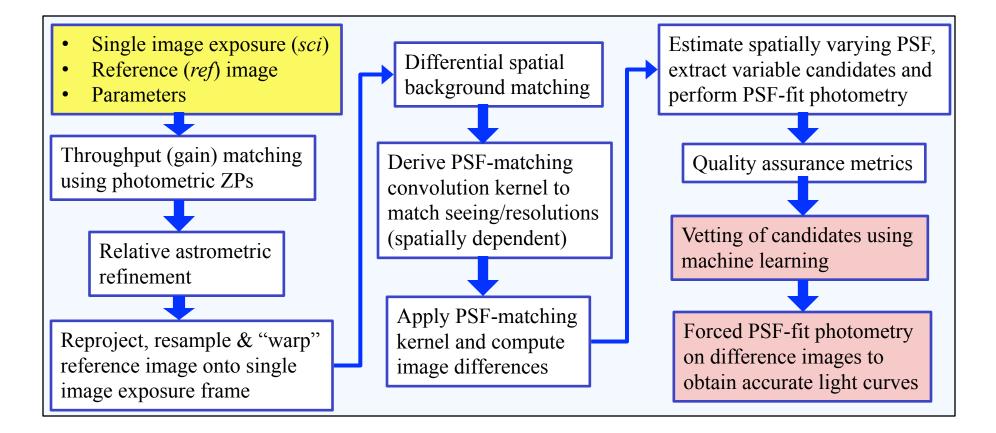
Edwin Hubble, 1949

PTF Data Flow and Processing



PTF Image-Differencing & Extraction pipeline (PTFIDE)

An engine for discovering candidate transients in near real-time for later vetting and follow-up



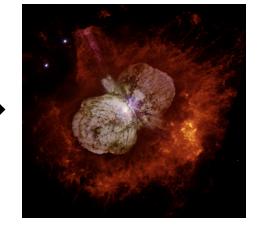
http://web.ipac.caltech.edu/staff/fmasci/home/miscscience/ptfide-v4.0.pdf

Validating PTFIDE on archival image data

- While validating pipeline, can we use new functionality to perform any new archival science?
- PLAN: look for activity at the locations of previously discovered supernovae before they exploded ⇒ missed pre-supernova eruptive mass-loss events
 - \Rightarrow tighter constraints on progenitors and massive-star evolution
- Started a campaign to reconstruct the light-curves of previously discovered supernovae by processing archived image data through new image-differencing pipeline
 - \Rightarrow with forced PSF-fit photometry, photometric accuracy is vastly improved
- Out of >10,000 PTF transient candidates, ~ 1900 are spectroscopically confirmed supernovae



Supergiant Star

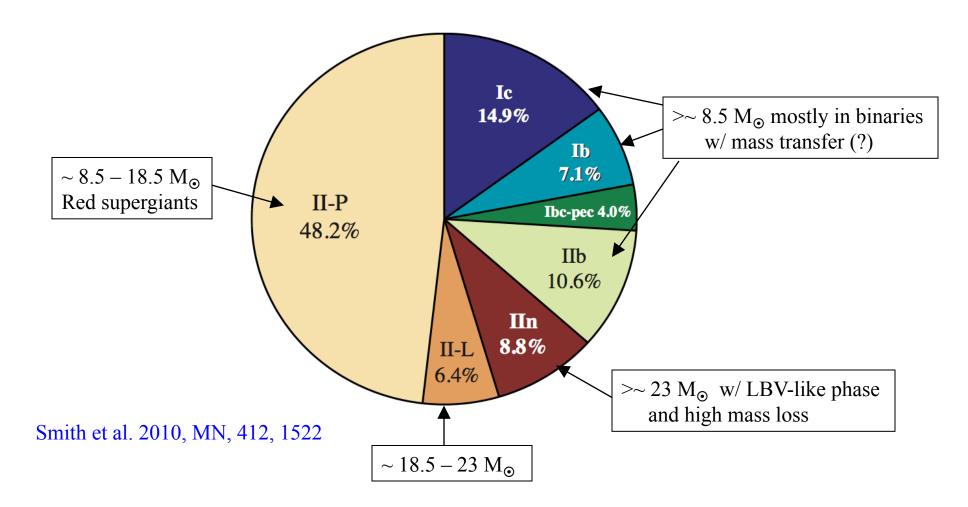


Eta Carina (Hubble)



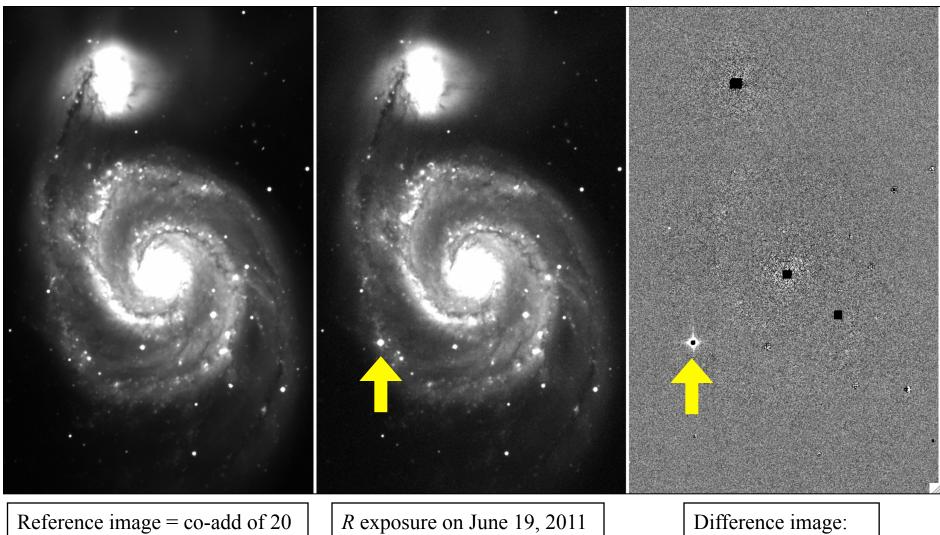
Type X Supernova?

Core-Collapse Supernova types and their progenitors



Degree of mass loss from progenitor star also determines which SN is seen

SN 2011dh (PTF11eon) in Messier 51

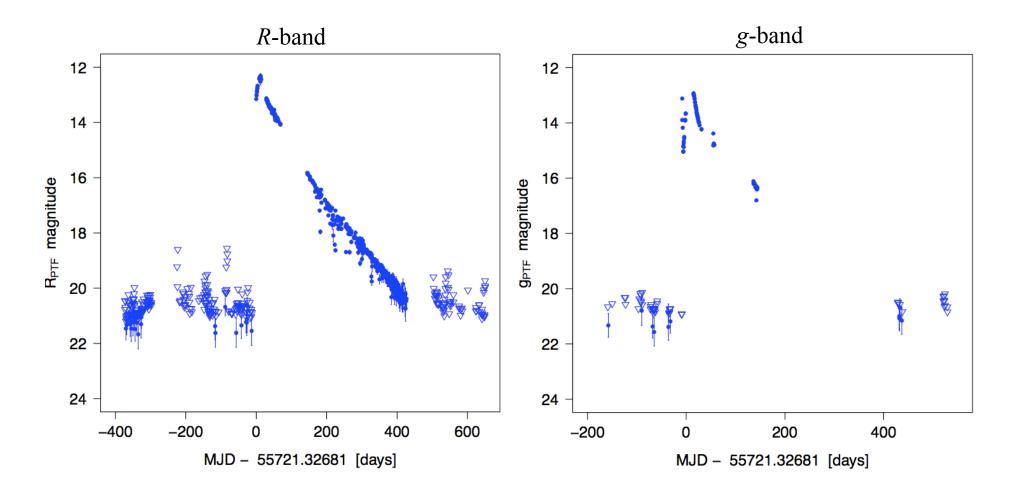


R exposures (pre-outburst)

R exposure on June 19, 2011Type IIb supernova ~ $10^9 L_{\odot}$

Difference image: exposure - reference

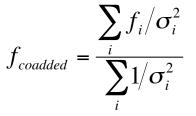
SN 2011dh light curves from PTF difference image photometry

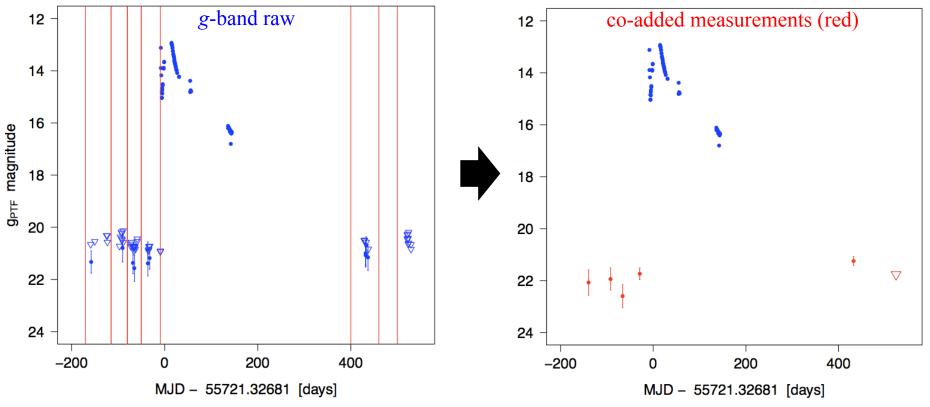


triangles: non-detections shown as 3.5σ upper limits

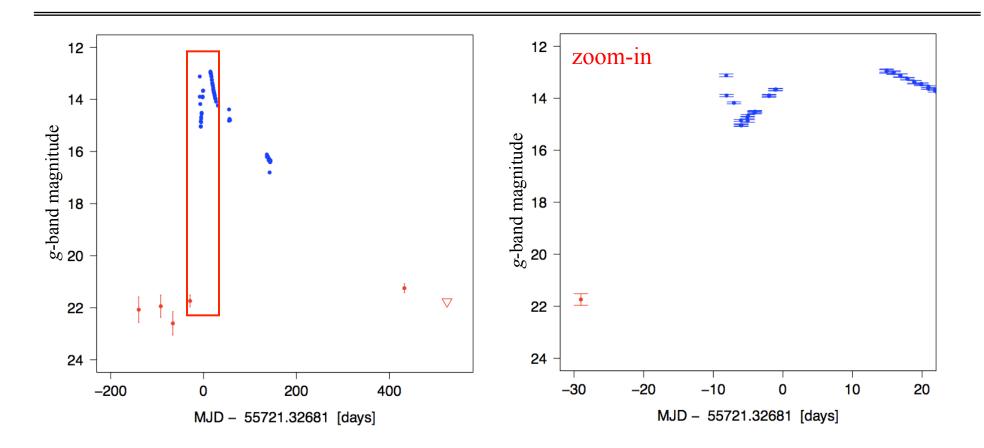
SN 2011dh light curves from windowed-averaging

- Combine measurements within windows to improve S/N or obtain tighter upper limits on non-detections
- **assumption:** fluxes in a window ~ constant with time. Can also collapse using more complex model based on prior (slope fit)





2011dh g-band light curve



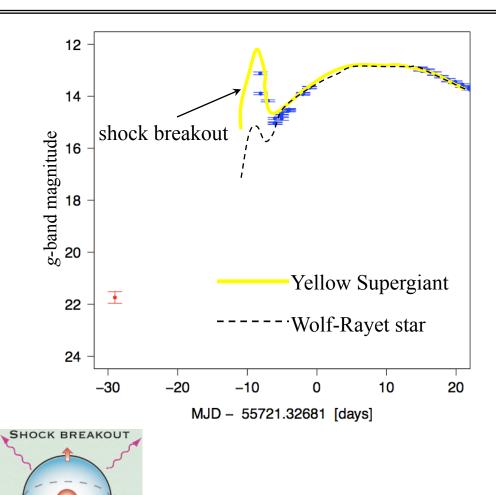
2011dh g-band light curve

- Bersten et al. (2012): light curve modeling: shape consistent with shock-breakout from photosphere of a yellow supergiant star.
- YSG progenitor consistent with pre-explosion HST photometry by Maund et al. (2011) & Van Dyke et al. (2011) with outer *H* envelope stripped, perhaps by a companion.

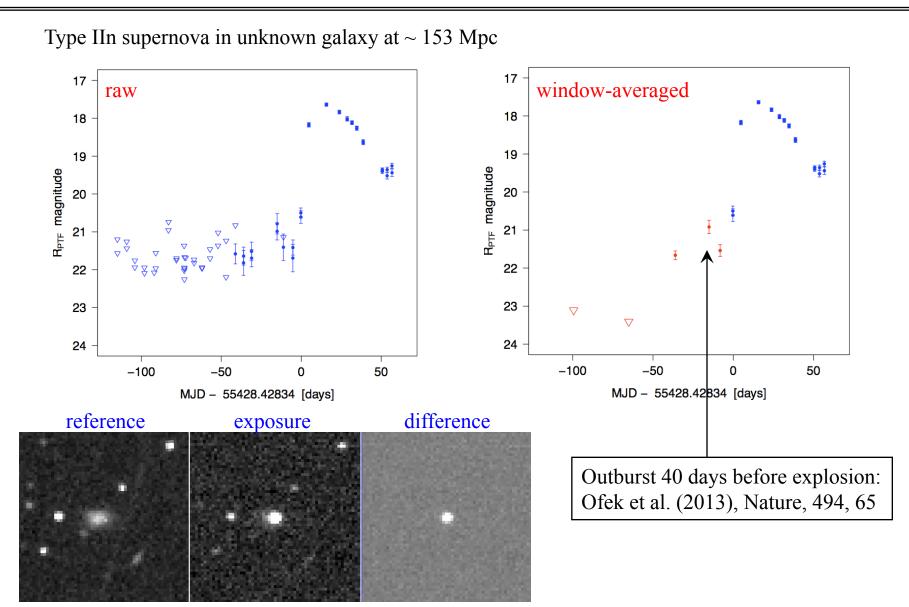
CORE COLLAPSE

📥 Fe 🖨

SHOCK PROPAGATION



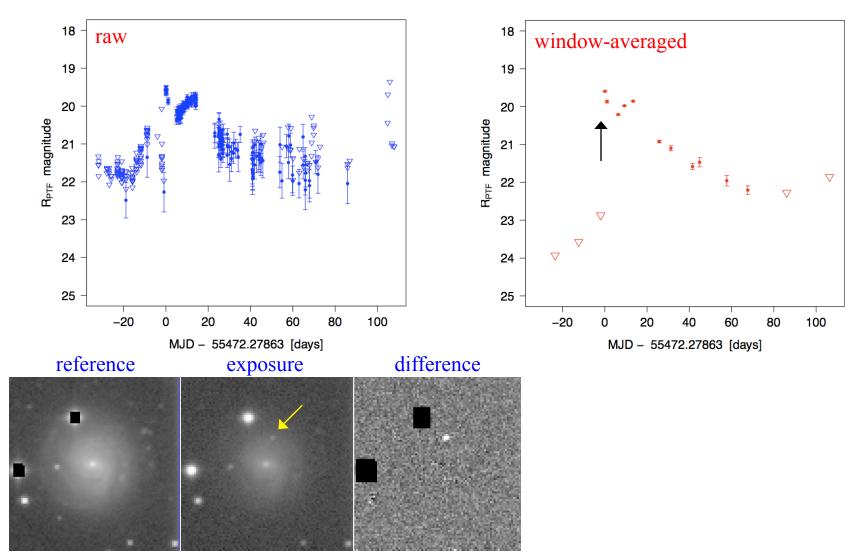
SN 2010mc (PTF10tel)



12

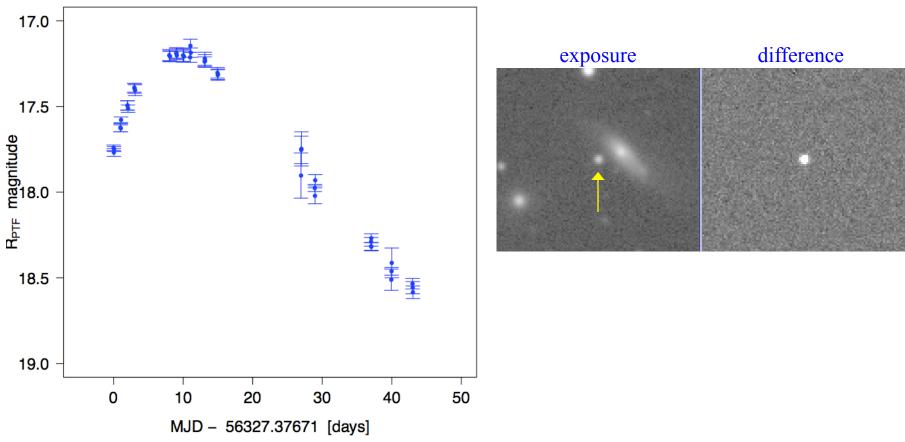
SN PTF10xfh

Type Ic supernova in NGC 717 at ~ 65 Mpc (Yi Cao, private communication)



SN PTF13ai (or PSN J12541585+0926259)

- Type Ia Supernova in galaxy PGC 43884 (~197 Mpc); discovered Feb 5, 2013
- One of the first to be discovered in new phase of PTF project: **iPTF**



Summary

- developed a suite of tools to enable accurate light curve generation from difference imaging
- good image subtraction with PSF-fit photometry and judicious time-window-averaging is the key at maximizing signal-to-noise!
- Started a campaign to reconstruct the light curves of confirmed supernovae using PTF archive
 - \rightarrow explore any missed pre-outburst (mass loss) episodes across different supernova types
 - \rightarrow better constraints on SN progenitors
 - \rightarrow better light-curve templates for classification or to support archival searches for missed SN