

ZTF Data System: Phase-II Plans & Schedule

Frank Masci & the IPAC/Caltech ZTF Team

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HEISING-SIMONS
FOUNDATION



Lawrence Livermore
National Laboratory



Trinity College Dublin
Coláiste na Tríonóide, Baile Átha Cliath
The University of Dublin

The ZTF Science Data System (ZSDS)

- The ZSDS is housed at IPAC, Caltech
- IPAC is a multi-mission science center (IRAS, 2MASS, *Spitzer*, WISE/NEOWISE, LSST, Euclid, WFIRST..)
- Responsibilities for ZTF:
 - data processing: from instrumental calibration to alert packets and lightcurves.
 - quality assurance; product archiving; user-interfaces to retrieve and analyze data.
 - management of data releases and user-support.
 - on-demand forced-photometry lightcurve service.

**Cahill Center
for Astronomy
& Astrophysics**



**Mt Palomar
~ 150 km**

IPAC

ZTF Computing & Storage at IPAC

- We ingest ~ 300 GB of raw camera data *per night* to generate ~ 3.8 TB in products.
- Processing is in “real-time” and alert packets leave IPAC ~ 10 to 20 minutes since observation.
- Expect ~ 5.7 PB in data products at end of Phase-II.
- Compute cluster consists of 66 compute nodes (1192 processor cores @2.5GHz each).

Racks containing 66 compute nodes



Archive fileserver/disk arrays



ZTF Public Release Data Products

- Like Phase-I, Phase-II public data releases will include:
 - Refinements to products and data quality flags in all previous releases (back to start of phase-I).
 - Products derived from additional observations:
 - Raw camera & calibration image data.
 - Calibrated single-epoch images, difference images, masks, PSFs, and source catalogs.
 - Lightcurves derived from positional re-matching across all epochs (back to start of Phase-I).
 - New reference images (co-adds), depth-of-coverage maps, and source catalogs.
 - Source database drawn from reference image catalogs to facilitate lightcurve retrieval.
- Public Alerts from the public surveys are not tied to any bulk data release
 - These continue to be distributed close to real-time for consumption by Alert Brokers.

Phase-II Public Data Release plan

- Move from a 6-month to 2-month release cycle for all data.
- **For public survey data:** move from a 6-month to 2-month embargo following DR4.
- **For private data** (partnership/Caltech science programs): continue with 18-month embargoing.

Release	Release Date	Public survey observation span	Private surveys observation span
DR4	12/09/20	03/17/18 – 06/30/20	03/17/18 – 06/30/19
DR5	03/31/21	03/17/18 – 01/31/21	03/17/18 – 09/30/19
DR6*	06/30/21	03/17/18 – 04/30/21	03/17/18 – 12/31/19
DR7*	08/31/21	03/17/18 – 06/30/21	03/17/18 – 02/29/20
DR8*	11/03/21	03/17/18 – 08/31/21	03/17/18 – 04/30/20
DR9*	01/05/22	03/17/18 – 10/31/21	03/17/18 – 06/30/20
...	...		

* Bimonthly release cycle

Data Access and Visualization Tools

<https://irsa.ipac.caltech.edu/Missions/ztf.html>

- Access is through IRSA at IPAC.
- Can search for images and source catalog files by position or object name (including SSOs), sources extracted from co-adds & their lightcurves; overlays, time series viewer w/ interactive manipulation.
- Accompanying APIs (command-line driven retrieval) are available for all services.

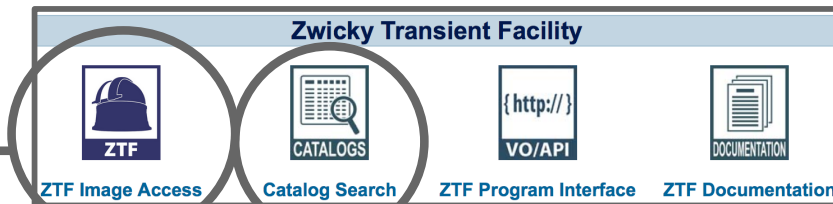


Image viewer and product retrieval

The screenshot shows the IRSA ZTF interface. At the top, there are tabs for "IRSA", "DATA SETS", "SEARCH", "TOOLS", and "HELP". Below the search bar, there is a table of search results. The table has columns for "ra", "dec", "field", "ccfid", "qid", "fid", "filtercode", and "obsdate". The first few rows of the table are:

ra	dec	field	ccfid	qid	fid	filtercode	obsdate
202.71695070000000	47.22087448000000	757	7	2	1	zg	2018-03-25 08:23:05+00
202.73263794000000	47.21753299000000	757	7	2	1	zg	2018-03-28 07:12:59+00
202.73615898000000	47.22155018000000	757	7	2	1	zg	2018-03-28 08:22:39+00
202.71972336000000	47.21063876000000	757	7	2	1	zg	2018-03-31 06:21:19+00
202.72316125000000	47.21021096000000	757	7	2	1	zg	2018-03-31 09:25:07+00

Below the table, there is an "Image Preview" section showing a grayscale image of a galaxy.

Lightcurve retrieval / Time Series Tool


The screenshot shows the "Lightcurve retrieval / Time Series Tool" interface. It has a "Column Selection" panel on the left with a table of columns: "magerr", "psfflux", "psffluxerr", "psfmag", "psfmagerr", and "catflag". The "Input Data" panel on the right shows a scatter plot of magnitude (mag) versus time (mjd (d)). The plot shows a series of data points with a clear periodic variation.

Period finding

The screenshot shows the "Period finding" interface. It has a "Periodogram" panel on the right showing a plot of power versus period (d). The plot shows a peak at a period of approximately 0.315 days. The "Set Period" panel on the left shows the selected period: "0.31496132".

Moving Object Search Tool (MOST)

asteroids pre-covered by ZTF imaging



NASA/IPAC INFRARED SCIENCE ARCHIVE

IRSA | DATA SETS | SEARCH | TOOLS | HELP Login

MOST - Moving Object Search Tool

The Moving Object Search Tool (MOST) can determine the orbit for a given solar system object then find images of object's predicted positions in select image datasets housed at IRSA (see [Instructions](#)). It can serve as a "precovery" if newly discovered objects were previously observed.

Image Dataset ztf

For ZTF: Time Range = 2017-10-15 to present
For complete range, leave limits blank (but this may take a long time)
Example: Antonia 2017-11-01 to 2017-12-31

Observation Begin (UTC) 2018-03-17	Observation End (UTC) 2020-10-22
Ephemeris Step Size (day) 1	Output Mode Regular

Create Fits and DS9 Region Files Tarballs

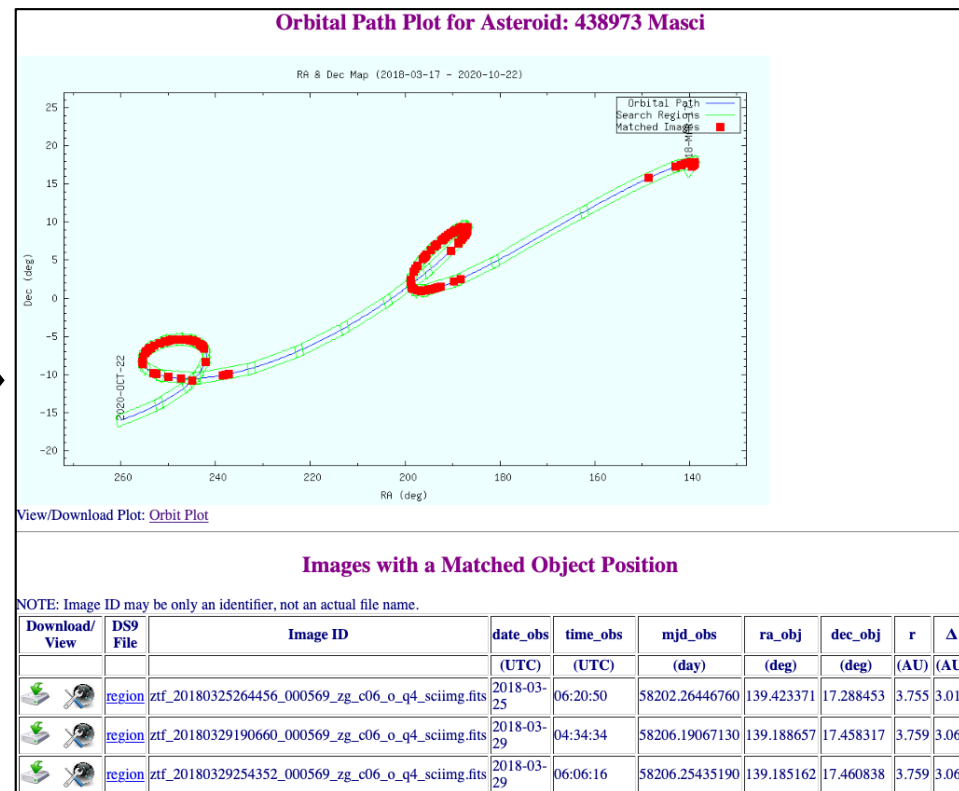
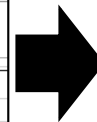
Solar System Object Name Input: Masci

Solar System Object NAIF ID Input:

MPC One-Line Element Input:

Object Type: Asteroid

Orbital Elements Manual Input:



Phase-II: major data-system upgrades

Capability / functionality	Release date
Public forced photometry service	2020-12-01
Forced photometry histories in alert packets (T – 30 days)	2020-12-20
Database, system, & infrastructure upgrades to support +3yr	2021-04-01
Bimonthly public release of file-based data products	2021-06-30 (commence DR6)
New lightcurve datastore format and database	2021-10-01
More frequent release of lightcurves (tied to new datastore)	2021-11-01 (commence DR8)
P60 archive & data access service	2021-10-01

Phase-II: other improvements & features

- Update astrometric calibration framework to use *Gaia Early DR3*.
- Update automated PSF-estimation software to further improve photometric accuracy: goal is $<\sim 1\%$ *absolute* precision (relative to PanSTARRS1), consistently on photometric nights.
- Propagate additional nearest *Gaia*-source metrics into alert packets, including proper motions.
- Include data in alert packets from the *PS1 Source Types & Redshifts with ML (PS1-STRM)* catalog.
- Refine Star/Galaxy classification scores of nearest PS1 sources in alert packets using latest ML.
- Reprocess subsets of improperly calibrated Phase-I data following upgrades.

Data Access & Documentation

- **Third Public Data Release: recipes for retrieving any ZTF data:**
<https://www.ztf.caltech.edu/page/dr3>
- **Access to Images, Catalogs, lightcurves, and analysis tools:**
<https://irsa.ipac.caltech.edu/Missions/ztf.html>
- **Science Data System Explanatory Supplement:**
https://irsa.ipac.caltech.edu/data/ZTF/docs/ztf_pipelines_deliverables.pdf
- **Science Data System paper:**
<https://iopscience.iop.org/article/10.1088/1538-3873/aae8ac>
- **Growing archive of raw public alerts and usage documentation:**
<https://ztf.uw.edu/alerts/public/>

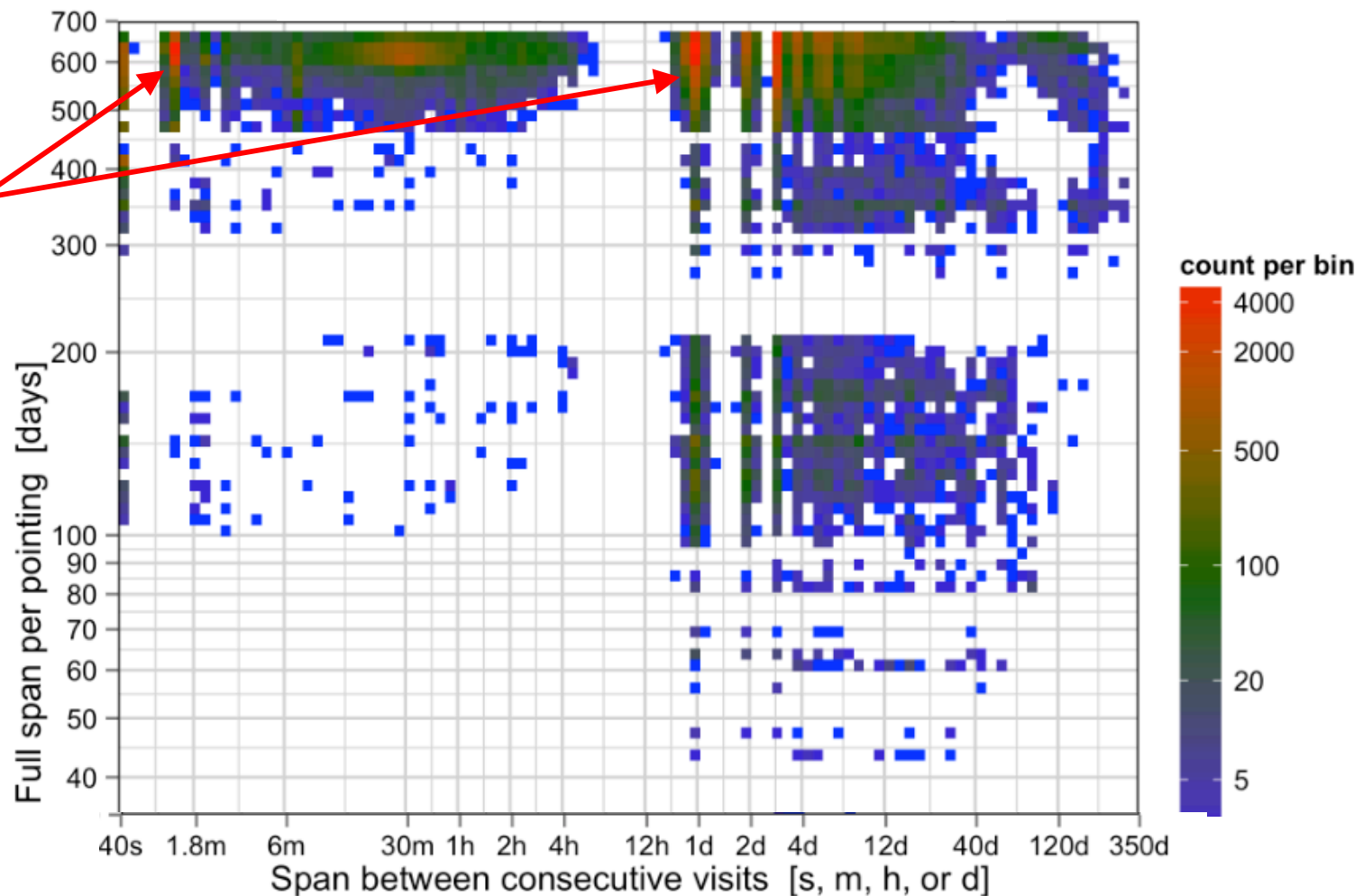
Back up slides

What can DR3 provide?

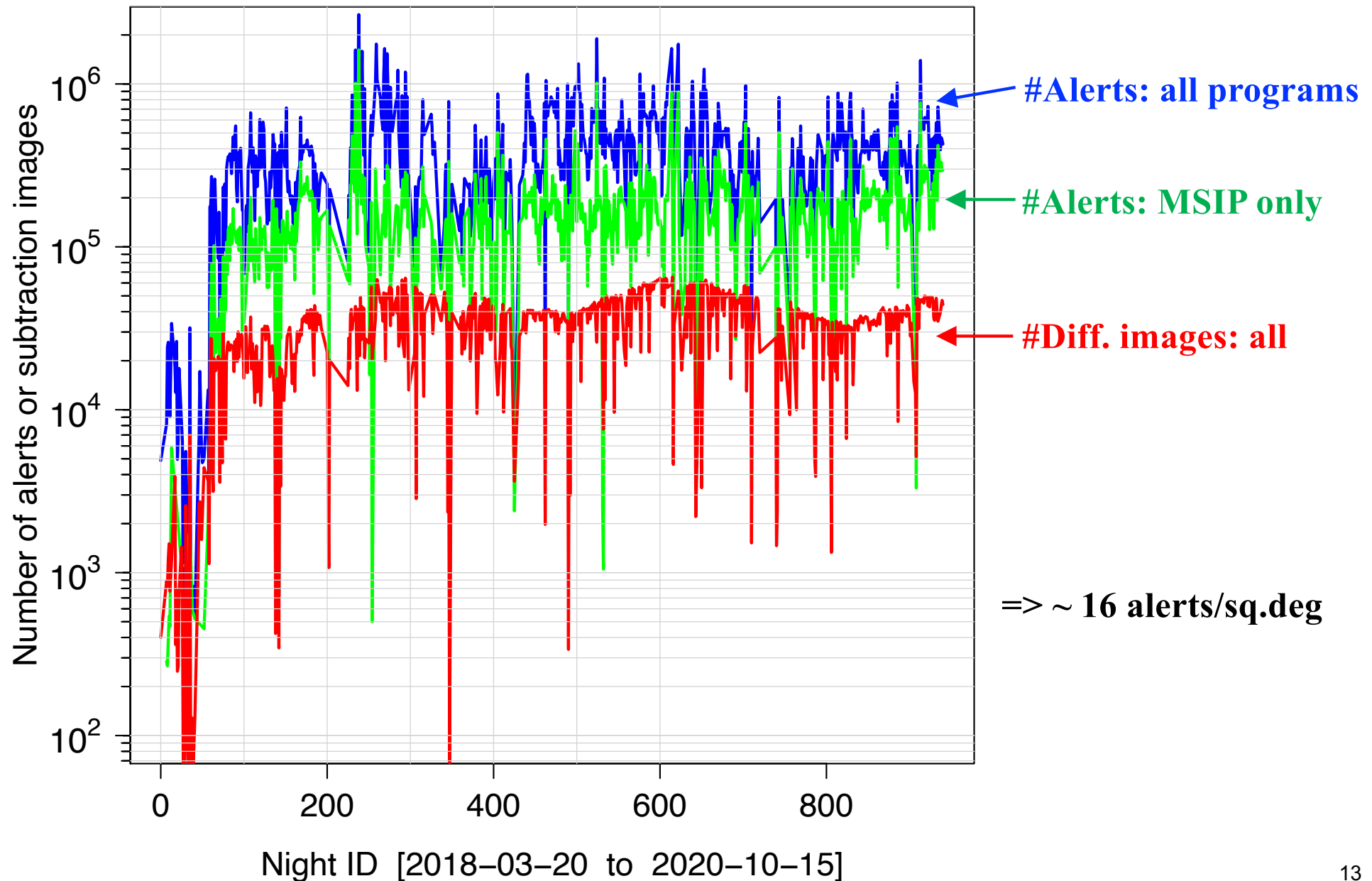
Lightcurve Timespans and Cadences

- Counts per bin are #visits *per 47 square degree field*.
- Sampling is non-uniform and irregular due to seasonal variation in field visibility from P48.
- Constrains the variability-related archival science projects possible with DR3.

includes galactic plane!

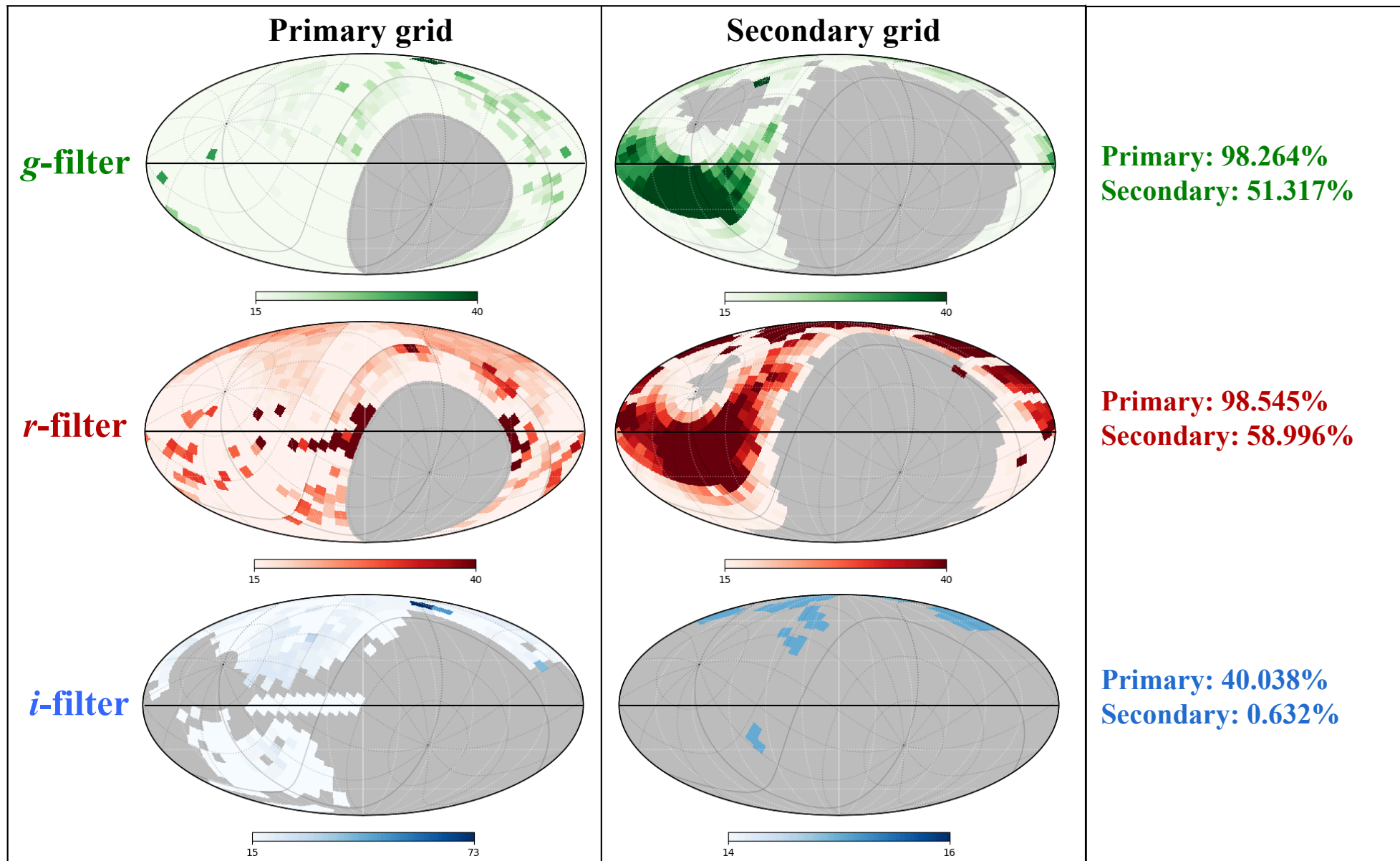


Alert Statistics (Mar 20, 2018 – Oct 15, 2020)



Reference Image Coverage as of Oct 15, 2020

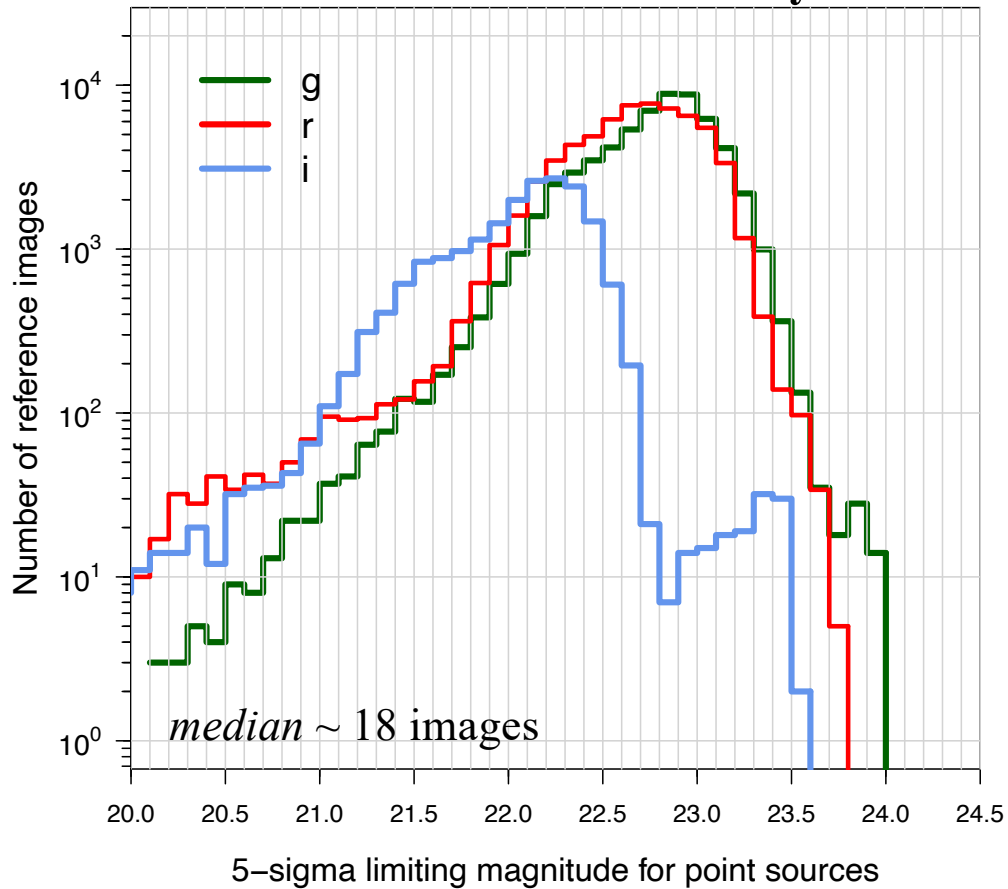
($l, b = 0, 0$ centered)



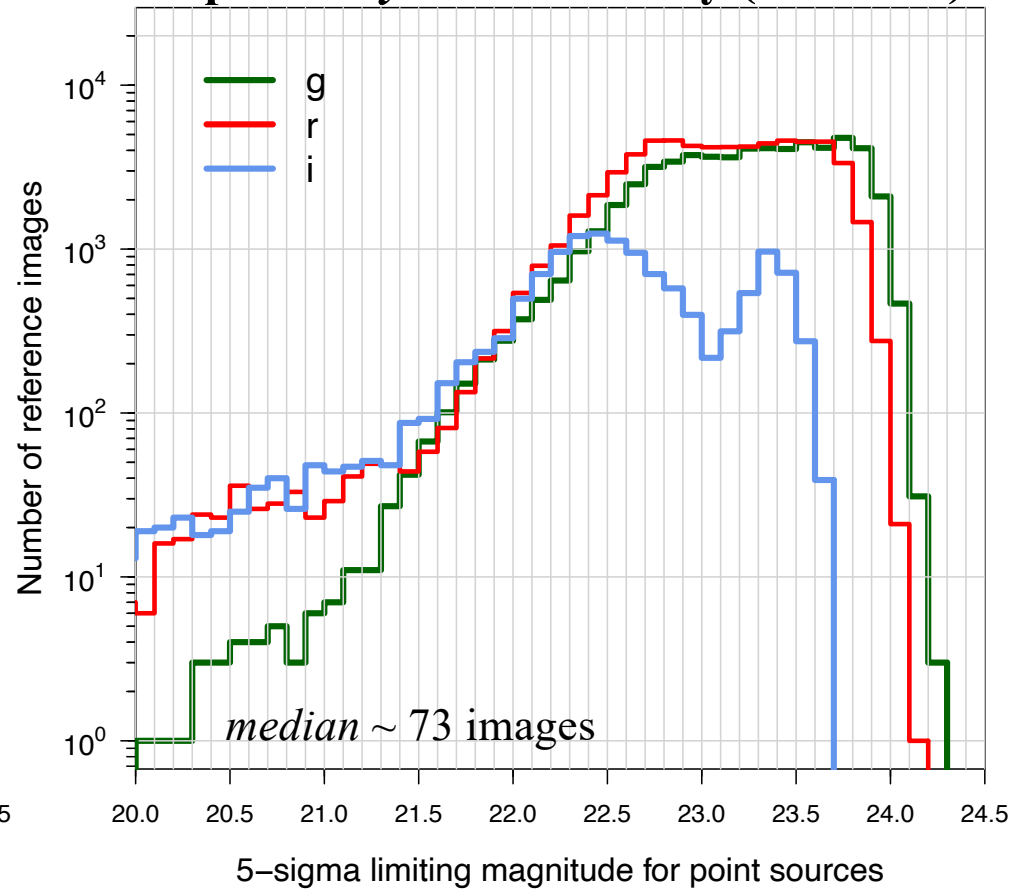
Reference Image Depths

Archived versus Special (*internal*)

Archived *Static* Library



Special *Dynamic* Library (for ToOs)



Fields with ultra-high source confusion:
noise & mag-limit estimators break down